

# Technical Implementation of SMOS Data in the ECMWF Integrated Forecasting System

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**Abstract**—The launch of the Soil Moisture and Ocean Salinity (SMOS) satellite of the European Space Agency opens the way to using a new type of satellite data that are very sensitive to soil moisture for numerical weather prediction. The European Centre for Medium-Range Weather Forecasts (ECMWF) has developed an operational chain which makes it possible to process SMOS data in near real time (NRT) and compare it with a model equivalent. This process has been very challenging. The main reasons are the particular characteristics of the SMOS observation system and the large volume of data. Despite these obstacles, SMOS data are being processed successfully in NRT within the ECMWF Integrated Forecasting System (IFS). The ultimate objective is to assimilate these data in the IFS. It is expected to have an impact on the weather forecast at short and medium ranges. Prior to assimilation experiments, the quality of the data has to be assessed. This can be done through monitoring activities. Monitoring is a routine task performed with all satellite data, and among other things, it makes it possible to localize temporal (or spatial) bias or drifts in the data, thus providing NRT reports to the calibration and validation teams, which can act accordingly. In this letter, the implementation of SMOS data in the ECMWF IFS for monitoring purposes is discussed. The system was developed using a simulated file for the NRT processor, and it was tested using real data from the first year since the launch date.

**Index Terms**—European Centre for Medium-Range Weather Forecasts (ECMWF), implementation, monitoring, Soil Moisture and Ocean Salinity (SMOS).

## I. INTRODUCTION

THE SUCCESSFUL launch of the Soil Moisture and Ocean Salinity (SMOS) satellite of the European Space Agency [1] is already providing an unprecedented new source of remotely sensed data that are sensitive to soil moisture over land and salinity over the oceans. Soil moisture has been extensively identified as a critical land variable due to its strong influence in the exchanges of water, energy, and carbon fluxes at the interface between the soil, vegetation, and the lowest level of the atmosphere [2]. A good estimation of soil moisture has a direct impact on precipitation and air temperature predictability at short and medium ranges [3], [4].

Passive L-band measurements are the most suitable ones for soil moisture retrievals. In this microwave region, attenuation from clouds and vegetation is smaller than that at higher frequencies [5]. Nonetheless, the heavy cost and technological

challenge of arranging a large antenna in L-band have prevented an earlier spatial L-band mission. For SMOS, an antenna of approximately 8 m in diameter would be necessary to meet the spatial resolution requirements of the mission [1], making the cost prohibitive with the current technology. In SMOS, this problem is overcome by applying the interferometric technique. Instead of one large antenna, 69 little receivers installed in three deployable Y-shaped arms of 3.5-m length collect the radiation emitted by the Earth's surface between 1400 and 1427 MHz. The phase difference measured between the individual receivers makes it possible to reconstruct an image which meets the science requirements, i.e., volumetric soil moisture with an accuracy of  $0.04 \text{ m}^3/\text{m}^3$  and a spatial resolution of 40–50 km [6]. As a numerical weather prediction (NWP) center, the European Centre for Medium-Range Weather Forecasts (ECMWF) is receiving a near-real-time (NRT) product, which is automatically recovered from the SMOS Data Processing Ground Segment. To fully take advantage of this product, ECMWF has implemented this new data type in the Integrated Forecasting System (IFS), which opens the possibility to monitor and assimilate the data within the IFS. This is a challenging task. In SMOS, the interferometric technique applied makes it possible to observe the same area under different views, thus providing multiangular and multipolarized observations of the same scene at different time stamps. Up to 150 records of brightness temperatures ( $T_B$ ) between  $0^\circ$  and  $65^\circ$  are provided per observed area. The angular resolution of the observations is very high. This measuring principle has the following two consequences: 1) the distribution of a unique data set with new features very different to any other source of satellite data used for NWP and 2) the production of a very large volume of data which cannot all be ingested in the IFS. Therefore, the data volume must be reduced significantly before a model equivalent is computed and compared with the observations. The previous characteristics of the SMOS observation system have raised great concern about the feasibility of integrating SMOS data in a complex NWP system. In this letter, the feasibility of its operational use is demonstrated. The main steps involved in the design of the SMOS data process chain are presented. The objective is to set up the structure necessary to operationally produce a comparison between a subgroup of SMOS observations and their model equivalent, as this is the main input for an assimilation scheme. The chain developed was tested with the first flow of available data and run successfully for one year.

## II. DATA PRODUCT USED AT ECMWF

The product used at ECMWF is the NRT which constitutes geographically sorted swath-based maps of  $T_B$ . The geolocated product received at ECMWF is arranged in an equal area grid

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95 system called ISEA 4H9 (Icosahedron Snyder Equal Area grid  
 96 with aperture 4 at resolution 9) [7]. For this grid, the centers of  
 97 the cell grids are at equal distances of 15 km over land, with  
 98 a standard deviation of 0.9 km. Over oceans, the grid has a  
 99 coarser resolution, which is half of the resolution over land, as  
 100 oceans are more homogeneous than continental surfaces. The  
 101 data are organized in messages. Each message corresponds to  
 102 a snapshot where the integration time is 1.2 s, as this is the  
 103 time in which all correlations of a single scene are measured.  
 104 On average, each message contains around 4800 observations  
 105 over land if the instrument runs in dual-polarization mode. In  
 106 this running mode, data set records are generated alternately  
 107 each 1.2 s at horizontal (HH) and vertical (VV) polarizations.  
 108 In full-polarization mode, all four Stokes parameters of the  
 109 radiation are collected by the antennas during four consecutive  
 110 integrations. Thus, the polarization state of the radiation is fully  
 111 described in this mode. Since the end of the commissioning  
 112 phase, the instrument on-board SMOS has been operating only  
 113 in full-polarization mode.

114

### III. IMPLEMENTATION

115 In this section, the main steps and challenges involved in the  
 116 implementation of SMOS data in the IFS are addressed. Before  
 117 SMOS data can be assimilated, the data go through a series of  
 118 tasks which have the objective of validating and comparing the  
 119 observations with a simulated value, thus producing an input  
 120 for the soil moisture assimilation scheme. These tasks can be  
 121 classified into the following two large groups:

- 122 1) data prescreening;
- 123 2) computations in the model grid.

#### 124 A. Data Prescreening

125 First, NRT raw data (see Section II) processed at the  
 126 European Space Astronomy Centre in Madrid (Spain) are re-  
 127 tried and slightly modified to feed the prescreening tasks.  
 128 These tasks perform quality control checks.

- 129 1) Generic checks: Files which do not contain crucial header  
 130 information are rejected. It is checked that date and time  
 131 are complete, geographic coordinates are not missing, and  
 132 instrument data correspond to SMOS data.
- 133 2) The validity of data is checked: Individual observations  
 134 are checked to be in a correct geographical position, and  
 135  $T_B$  is checked to be in the range of physically reasonable  
 136 values (not lower than 50 K and not greater than 350 K),  
 137 which is also a practical hard radio-frequency interfer-  
 138 ence (RFI) filter.
- 139 3) Data are thinned to reduce the volume of SMOS data  
 140 processed within the IFS.

141 Data thinning is a critical step insofar as it selects not  
 142 only which data from the original files will be monitored but  
 143 also which data could be used to correct the soil moisture  
 144 state through an assimilation experiment. The daily volume of  
 145 SMOS data arriving in ECMWF archives in NRT is of about  
 146 8 GB, which is by far one of the greatest sources of satellite data  
 147 received at ECMWF. This amount of data cannot be introduced  
 148 in the IFS for just one single satellite instrument, taking into  
 149 account that data from many other satellites are used simulta-  
 150 neously. For SMOS, only 5%–10% of the initial data volume

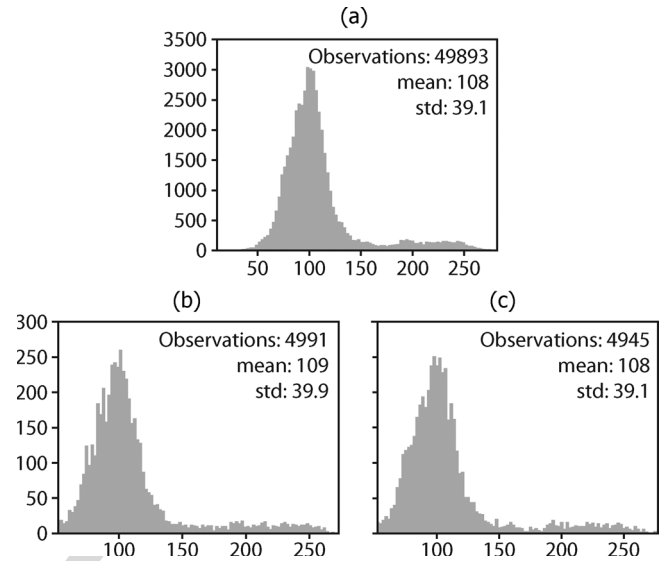


Fig. 1. Histograms of brightness temperatures for an SMOS-simulated file for December 17, 2010. (a) Histogram of  $T_B$  for testing file. (b) Histogram of  $T_B$  for EXP1. (c) Histogram of  $T_B$  for EXP2.

can realistically be ingested in the IFS. Thinning is therefore a  
 151 mandatory step and also prevents redundant observations in the  
 152 assimilation system. 153

Data can be thinned in many different ways. For monitoring  
 154 purposes, the thinned subset of observations should be repre-  
 155 sentative and keep the same statistical characteristics as the  
 156 original data set. In this letter, two simple experiments are  
 157 shown in order to investigate the characteristics of a thinned  
 158 subset of SMOS observations. In the first experiment (EXP1),  
 159 only observations containing integer incidence angles  $\pm 0.1^\circ$   
 160 are selected. This offset parameter was tuned in order to filter  
 161 approximately only 10% of the initial number of observations.  
 162 The rationale for this filter is that, even though the surface  
 163 emission is not a linear process with the incidence angle, the  
 164 land microwave emission should not be very sensitive to small  
 165 differences in the angular illumination of the target. Thus, a  
 166 significant data reduction should be achieved while keeping the  
 167 original angular signature of the observations. In the second  
 168 experiment (EXP2), a simple filter which keeps only one out of  
 169 ten observations from a snapshot was applied. This is a practical  
 170 way to keep, roughly, only 10% of the original data volume,  
 171 making it compatible with the IFS memory capabilities. 172

With a view to implementing a thinning filter in the IFS, these  
 173 two experiments were tested with an SMOS-simulated NRT file  
 174 for December 17, 2010. That was the only data available. This  
 175 file contains several snapshots covering mostly the Arabian Sea  
 176 but also some land in Oman and Saudi Arabia. In total, it has a  
 177 volume of 1 067 724 B, corresponding to 48 993 observations.  
 178 The histogram of simulated  $T_B$  is shown in Fig. 1(a). It clearly  
 179 shows the sea (predominantly) and land features. The mean is  
 180 108 K, whereas the standard deviation is 39.1 K. Fig. 1(b) and  
 181 (c) shows the histograms of  $T_B$  for EXP1 and EXP2, respec-  
 182 tively. In EXP1, exactly 10% of the initial data set was selected,  
 183 whereas in EXP2, 9.91% was selected. For EXP2, the statistics  
 184 remain as in the original data set, whereas in EXP1, they are  
 185 slightly degraded. The main advantage of EXP2 is that all  
 186 incidence angles are still available for monitoring purposes and  
 187 it allows greater flexibility for further processing. For example, 188

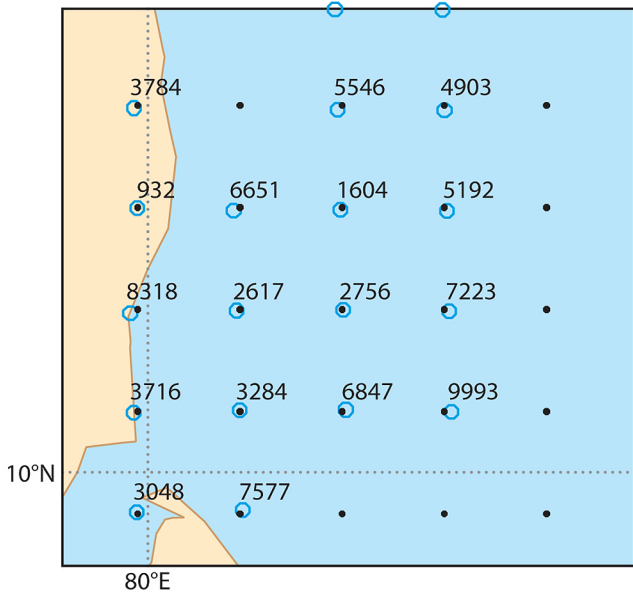


Fig. 2. (Blue circles) Nearest SMOS observations to the (black dots) ECMWF T159 model grid points. The values are the distances between grid points and the nearest SMOS observation, in meters.

189 by using EXP2, data can be averaged in angular bins at later  
 190 steps. Furthermore, the implementation is very simple, and only  
 191 one parameter is needed to control this filter. For EXP1, the  
 192 offset parameter has to be optimized each time to control the  
 193 number of observations filtered, and not all angular geometries  
 194 can be monitored. Thus, the initial implementation of a filter  
 195 type as in EXP2 was preferred for monitoring purposes. Due  
 196 to operational constraints, this process has to be completed  
 197 very quickly. On average, 12 h of data are processed in almost  
 198 1 min using eight processors in parallel. This makes SMOS  
 199 prescreening tasks fully compatible with the current ECMWF  
 200 operational system.

201 *B. Computations in the Model Grid*

202 The implementation of SMOS data monitoring in the IFS was  
 203 carried out in the model grid. This has several advantages.

- 204 1) All the background fields necessary to simulate  $T_B$  at  
 205 the top of the atmosphere are computed and available in  
 206 the model grid point space. Thus, it avoids interpolating  
 207 physical quantities to an observation location.
- 208 2) Other satellite data that are sensitive to soil moisture,  
 209 such as AMSR-E data in C-band, are also available in  
 210 the model space, making a comparison with other satellite  
 211 data possible. The subset of observations which are  
 212 selected after the prescreening filters undergoes (roughly)  
 213 a two-step process.
- 214 a) Observations are brought into the model grid point space  
 215 at the required model resolution by using the nearest  
 216 neighbor technique. At the same time, a mask of the  
 217 flags containing information of the grid point is created.  
 218 Fig. 2 shows the SMOS observations for the simulated  
 219 file on December 17, 2010, collocated to the nearest  
 220 grid point. For the sake of clarity, a magnified area over  
 221 the south Indian coast is shown, and a rather coarse  
 222 grid T159 is used ( $\sim 125$  km). The distance limit beyond

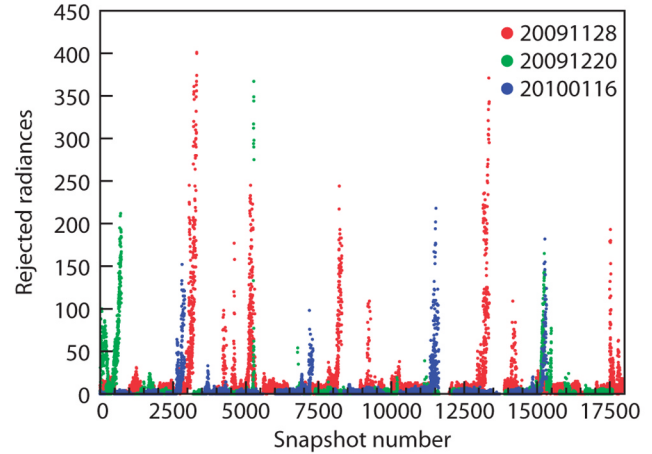


Fig. 3. Number of rejected observations in the prescreening tasks for 6 h of NRT SMOS data.

which observations are rejected was set to 10 000 m. The 223  
 224 number of observations monitored depends on the model 224  
 225 grid resolution and the distance limit parameter. At T799 225  
 226 ( $\sim 25$  km) or higher spectral resolution, SMOS observa- 226  
 227 tions within the distance limit were found for all grid 227  
 228 points (not shown).

- b) For each observation flagged as being the closest to a 229  
 230 model grid point, a background value is simulated with 230  
 231 an L-band forward model operator. Thus, the innovation 231  
 232 vector is obtained as the main input for a soil moisture 232  
 233 analysis. The forward model interfaced to the IFS is 233  
 234 the Community Microwave Emission Model [8], [9], as 234  
 235 explained in [10].

IV. RESULTS WITH REAL DATA

The operational chain described in Section III was tested 237  
 238 with real  $T_B$  data from the commissioning phase. No distinction 238  
 239 between ascendant and descendent orbits was made. These data 239  
 240 sets were the following:

- Data set 1) November 28, 2009; 241
- Data set 2) December 15, 2009; 242
- Data set 3) December 20, 2009; 243
- Data set 4) January 16, 2010. 244

A. Prescreening Results and Model Equivalent

The quality checks listed in Section III-A were tested with 246  
 247 data sets 1), 3), and 4). Fig. 3 shows the number of individual 247  
 248  $T_B$  values rejected as a function of the first 18 000 snapshots. 248  
 249 This corresponds to the first 6 h of collected data for these 249  
 250 days. During the commissioning phase, the SMOS operational 250  
 251 mode was alternated between dual and full polarizations so as 251  
 252 to test and select the optimal mode for the instrument. Data 252  
 253 sets tested in this section contain data in both modes. Thus, 253  
 254 for comparison purposes, only snapshots with fewer than 5000 254  
 255 subsets were used because they correspond to pure HH- or 255  
 256 VV-polarization integrations. This figure clearly shows that the 256  
 257 number of rejected radiances is the greatest for November 28, 257  
 258 when the data were not yet calibrated, whereas the number is 258  
 259 significantly reduced for December 20 and January 16. Table I 259  
 260 shows a quantitative comparison between the three data sets. It 260



TABLE I  
NUMBER OF OBSERVATIONS REJECTED FOR 6 h OF DATA DURING  
THE EARLY QUALITY CHECK PHASE

Date	snapshots	subsets	rejections	% rejected
28-11-2009	17940	28203176	147185	0.52
20-12-2009	17592	28739029	58967	0.21
16-01-2010	15347	24322415	34386	0.14

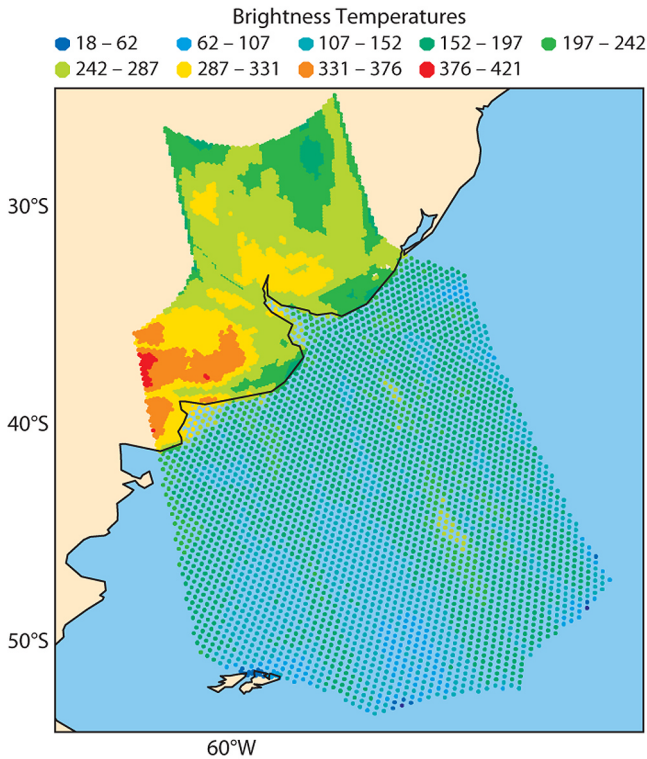


Fig. 4. Brightness temperatures over the east of Argentina (in kelvins), in VV polarization for an NRT file received December 15, 2009.

TABLE II  
THINNING FILTERS EXP1 AND EXP2. “Nbr OBS” IS THE NUMBER OF  
OBSERVATIONS, “MEAN” IS THE BRIGHTNESS TEMPERATURE MEAN  
VALUE (IN KELVINS), AND “STD” IS THE STANDARD DEVIATION

	Original file	after EXP1	after EXP2
Nbr obs	522566	52589	52306
mean	189	191	189
STD	63.7	63.1	63.5

261 shows how the quality of the data is the best in January 2010, 262 with only 0.14% of observations rejected after the first group of 263 quality checks. Although this percentage is small in November, 264 it is relatively more significant than those in December 2009 265 and January 2010.

266 Data set 2) was used to test the thinning filters described as 267 EXP1 and EXP2 in Section III-A. In total, this file contains 268 only 211 snapshots (southeast Argentinian coast) with 537 132 269 observations (Fig. 4). The results are summarized in Table II 270 for the VV polarization. The results for HH polarization were 271 equivalent. In both cases, the number of observations filtered 272 was very similar. For EXP2, the statistical signature of the 273 thinned subset is closer to that of the original data set, but 274 more importantly, there is still the possibility to monitor any 275 incidence angle. With this single filter, on average, ten angular 276 observations are considered for each grid point. In contrast to an 277 EXP1-type filter, this filter approach has better control over the

278 volume of the filtered data set. By using a series of processors in 279 parallel, the collocation to the model grid is performed quickly 280 and efficiently. At this stage, each model grid point is associated 281 with the closest observation, as long as it is within the distance 282 limit parameter. Then, for each grid point, the geometry of the 283 closest observation is used to compute a model equivalent and 284 compare it with the observation. At the end of this process, each 285 grid point is associated with at least one model equivalent.

## B. Continuous Monitoring of SMOS Observations

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287 The feasibility of using the previous chain operationally 288 was tested with several months of real data. SMOS data were 289 monitored by systematically producing global maps of NRT  $T_B$  290 for different incidence angles (0, 10, 20, 30, 40, 50, and 60) 291 and for both HH- and VV-polarization modes. Plots for 2010 292 are available at [http://www.ecmwf.int/research/ESA\\_projects/](http://www.ecmwf.int/research/ESA_projects/SMOS/monitoring/2010/2010.html) 293 SMOS/monitoring/2010/2010.html. A simple inspection of 294 these figures makes it possible to observe not only the angular 295 evolution of the data for each polarization state but also a 296 significant improvement in the quality of the observations, 297 specially after the instrument was calibrated. Fig. 5 shows  $T_B$  298 at 40° incidence angle and VV-polarization state for days 1 299 (top), 3 (middle), and 4 (bottom). It shows a clear evolution 300 in the quality of the data, from day 1 (top) to day 4 (bottom). 301 The day in November (top) is shown to be very noisy. These 302 were data received within the first two weeks of the instrument 303 “Switch-On Phase,” which obviously had not benefited yet 304 from a good calibration. In December, a major calibration event 305 took place, and the difference in the product is quite significant 306 when comparing the top and middle figures. Improvements 307 are present almost everywhere. The data are even better for 308 January 16, although this needs a closer examination and quan- 309 titative results to confirm it (see Table I). Results for the HH 310 polarization were equivalent. As SMOS is a research mission, it 311 was also important to check the correct functioning of the novel 312 instrument. Hence, it was checked that the  $T_B$  values got colder 313 with increasing the incidence angle for HH polarization, and the 314 opposite behavior was shown for the VV polarization (see, for 315 example, the online plots), with both displaying values within 316 an acceptable physical range. It is also an objective of data 317 monitoring activities to report on possible spatial or temporal 318 effects on the data: The most outer sides of the satellite track 319 look colder than the inner part, which is due to the extended 320 alias-free field of view, of lower quality than data closer to the 321 center of the track. There is still residual RFI over Europe, the 322 Middle East, and Asia, which is particularly straightforward to 323 spot when the data look very “red” and noisy. The data are of 324 relatively better quality over the whole of America, Australia, 325 and southern Africa.

## V. SUMMARY

326

327 Integrating a new type of satellite data with innovative fea- 328 tures in a complex NWP system and making it fully compatible 329 with an NRT structure is a very challenging task. This is the 330 case for the SMOS NRT product arriving at ECMWF. The 331 large volume of SMOS data and the large number of angular 332 views per pixel have raised concerns about the feasibility of 333 using this product in an operational NWP context. Despite these 334 challenging features, the operational use and the feasibility of

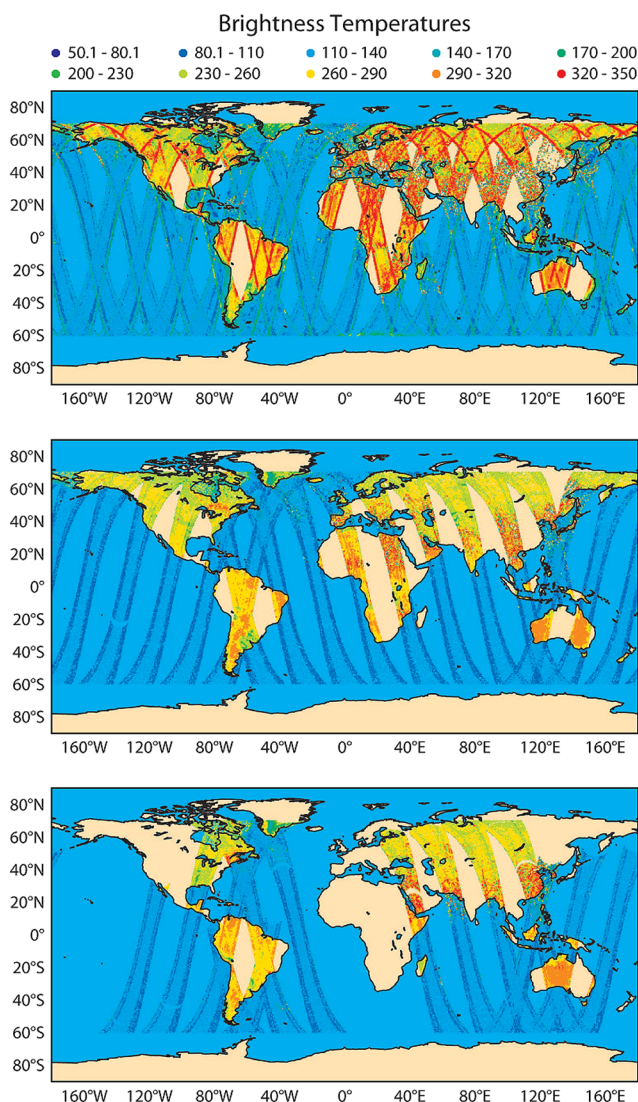


Fig. 5. Brightness temperatures for NRT SMOS data at 40° incidence angle and VV-polarization state. The figure on the top is for November 28, 2009, the middle figure corresponds to December 20, 2009, and the bottom figure corresponds to January 16, 2010.

335 processing a large amount of SMOS data in NRT were shown.  
 336 The objective was to set up the structure required to compare  
 337 SMOS measurements with a model equivalent. This opens the  
 338 possibility of investigating the benefits that these observations  
 339 can bring for weather forecast skill. Thus, rather than providing  
 340 an exhaustive analysis of the observations, this letter is more  
 341 oriented toward demonstrating that the SMOS data monitoring  
 342 chain is efficient and compatible with an NRT system.  
 343 The operational chain proposed in this letter is able to  
 344 cope with more than 8 GB of SMOS data a day, i.e., more  
 345 than 300 million observations. Each observation is subjected  
 346 to an exhaustive quality control; then, the whole data set is  
 347 strongly thinned. Although the thinning filter proposed is quite  
 348 simple (keeping only one out of ten observations), it keeps all  
 349 the incidence angles (for monitoring purposes) and has good  
 350 control over the size of the filtered data set. However, the  
 351 optimal thinning filter to be applied to the data is not obvious.  
 352 For example, possible approaches may include averaging the  
 353 data within predefined angular bins, discarding snapshots where

a grid point has a  $T_B$  value exceeding a given threshold, or  
 removing noisy data at low incidence angles. Certainly, the  
 development of an enhanced thinning filter deserves further  
 investigation, and it is a key activity as assimilation experiments  
 may depend strongly on the selected subsample of data.

Forward modeling on the model grid (using the geometry of  
 the nearest SMOS observation) proved to be efficient and fast  
 if several processors were used simultaneously. The analysis  
 of the first batch of real data using this structure proved to be  
 useful. It suggested a clear enhancement in the data quality  
 during the first months of the mission, both qualitatively and  
 quantitatively. There are still strong sources of RFI remaining  
 in Europe and Asia, whereas a visible data degradation is  
 still observed at the edges of the satellite track over oceans  
 at high incidence angles. The systematic production of these  
 plots in NRT is an excellent way to monitor the data just a  
 few hours after sensing time and to quickly inform calibration  
 and validation teams about trends or drifts in the data. In this  
 context, SMOS data monitoring will be supported through the  
 provision of the statistics of the modeled values in the grid  
 point space, where first-guess departures are computed. These  
 statistics are currently being obtained for several weeks of data,  
 and their analysis will be presented in a follow-up paper totally  
 devoted to this aspect.

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## AUTHOR QUERIES

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AQ1 = This sentence was reworded for clarity. Please check if the intended meaning was retained.

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